

The background of the slide is a composite image. The upper half shows a dark, starry space with a large, detailed image of the Moon in the upper left corner. The lower half shows a curved horizon of the Earth, with a bright blue sky and white, swirling clouds.

# Predicting Material Performance in the Space Environment from Laboratory Test Data, Static Design Environments, and Space Weather Models

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# Introduction

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- **Materials are evaluated for use in space environments by laboratory exposure to VUV/UV, AO, and charge particle environments to determine effects on material properties**
- **Standard “static” design environments are typically used to establish exposure periods and the corresponding photon, AO, and charged particle fluence to meet mission requirements**
- **Questions:**
  - How well do static models represent the real environment?
  - What is the contribution of “space weather” events to material exposure environments?
- **Today’s presentation will**
  - Examine VUV/UV environments used in laboratory tests with emphasis on surface exposures
  - Examine importance of “space weather” event contributions to environment

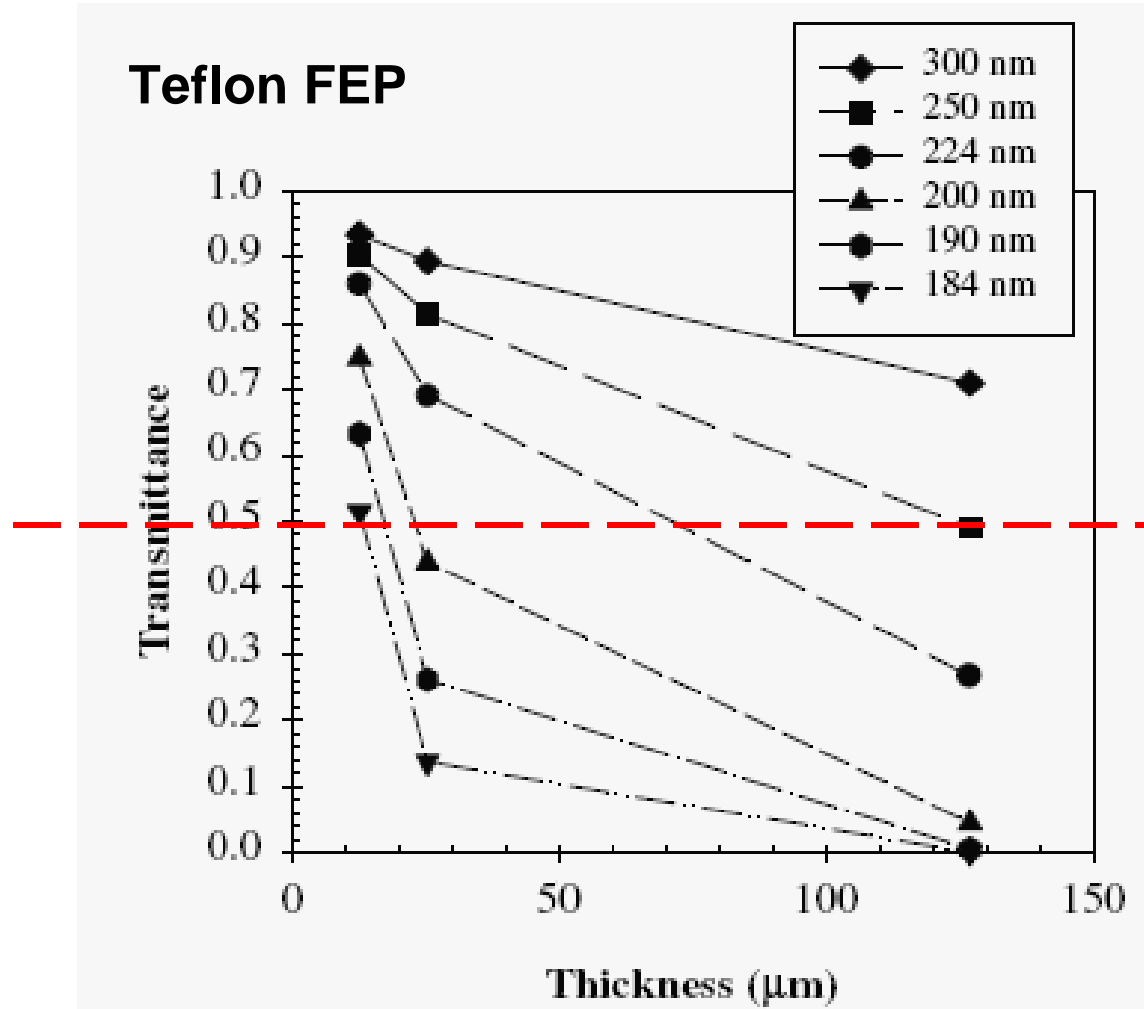


# VUV/UV Penetration Depth

- Long wavelengths penetrate deeper into polymers

| $\lambda$ (nm) | 50% depth ( $\mu\text{m}$ ) |
|----------------|-----------------------------|
| 300            | ----                        |
| 250            | 128                         |
| 224            | 65                          |
| 200            | 22                          |
| 190            | 17                          |
| 184            | 14                          |

- VUV wavelengths where solar variability is strongest primarily impacts material surfaces

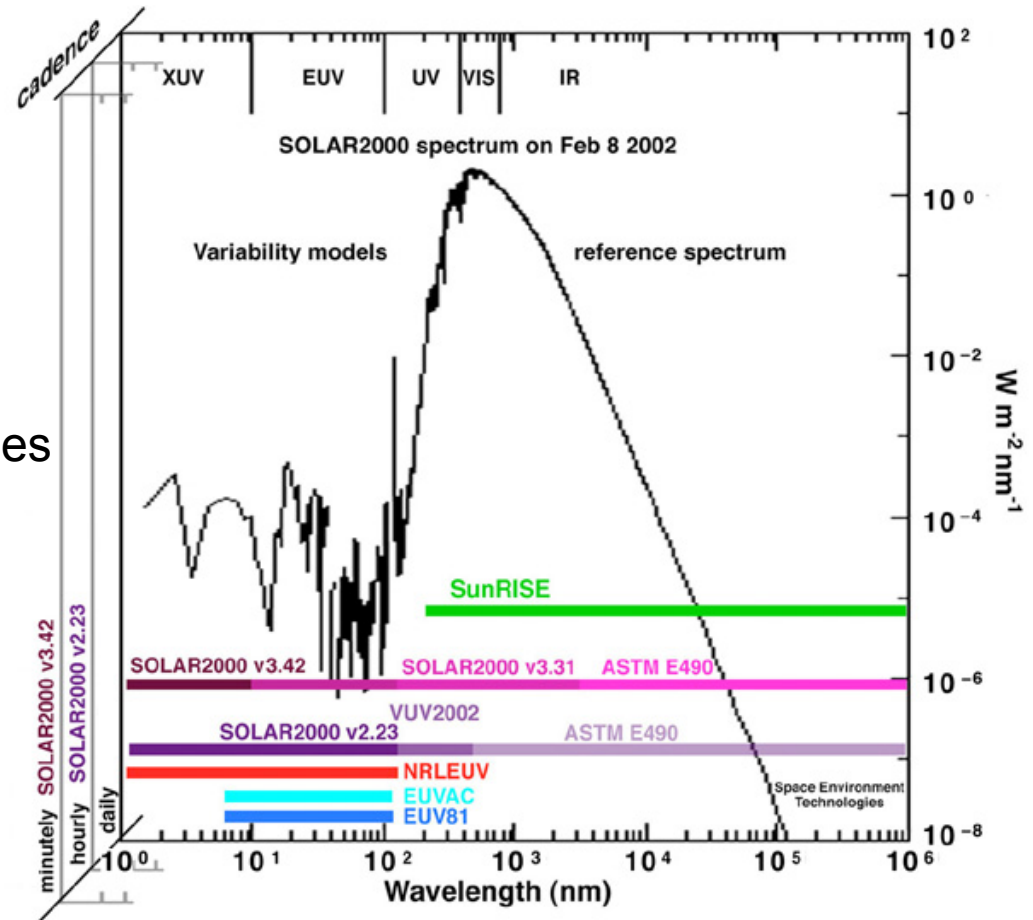


[Dever et al., 2002]



# Solar Spectrum and Models

- Static models
  - NRLEUV
  - EUVAC
  - EUV81
  - ASTM E490
- Climatology
  - Solar2000 (S2K)
  - Space Environment Technologies  
Tobiska et al.  
constant  $122.5 - 1 \times 10^6$  nm  
variable  $0.5 - 121.5$  nm  
 $\Delta t = 1$  day
- Space weather
  - Flare Irradiance Spectral Model
  - LASP/CU Boulder  
Chamberlin et al.  
 $0.1 - 194$  nm  
 $\Delta t = 1$  minute



[Tobiska and Nusinov , 2004]



# Solar XUV/EUV/UV Variability

## Solar2000

- Static ASTM E490
- Variable XUV/EUV

## Reference Spectrum

- Mean S2K
- Nomenclature

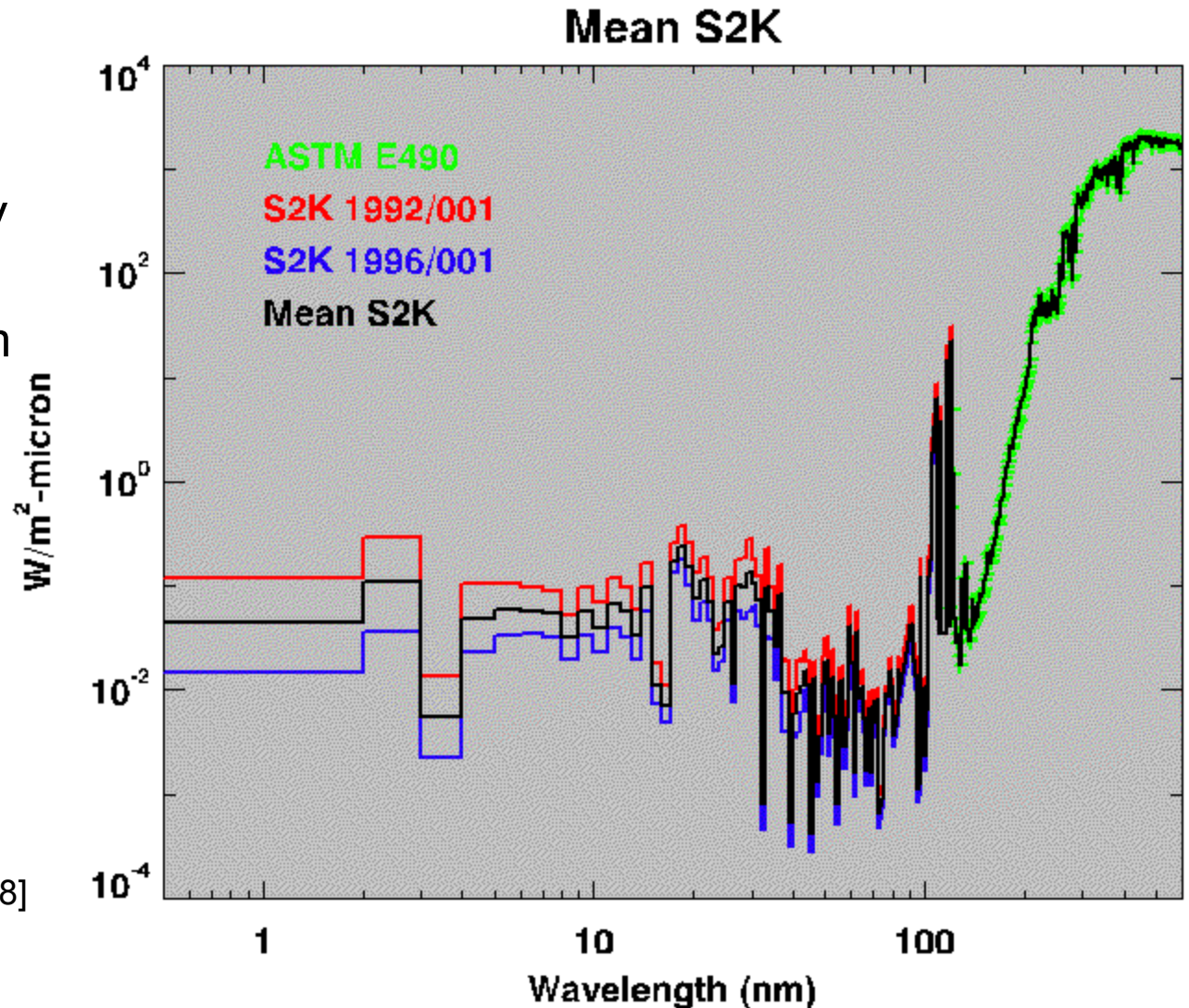
XUV  $0.1 \leq \lambda < 10$

EUV  $10 \leq \lambda < 200$

UV  $200 \leq \lambda < 400$

VUV  $20 \leq \lambda < 200$

[ISO\_DIS\_21348\_E\_revB, 2008]





# Solar XUV/EUV/UV Variability

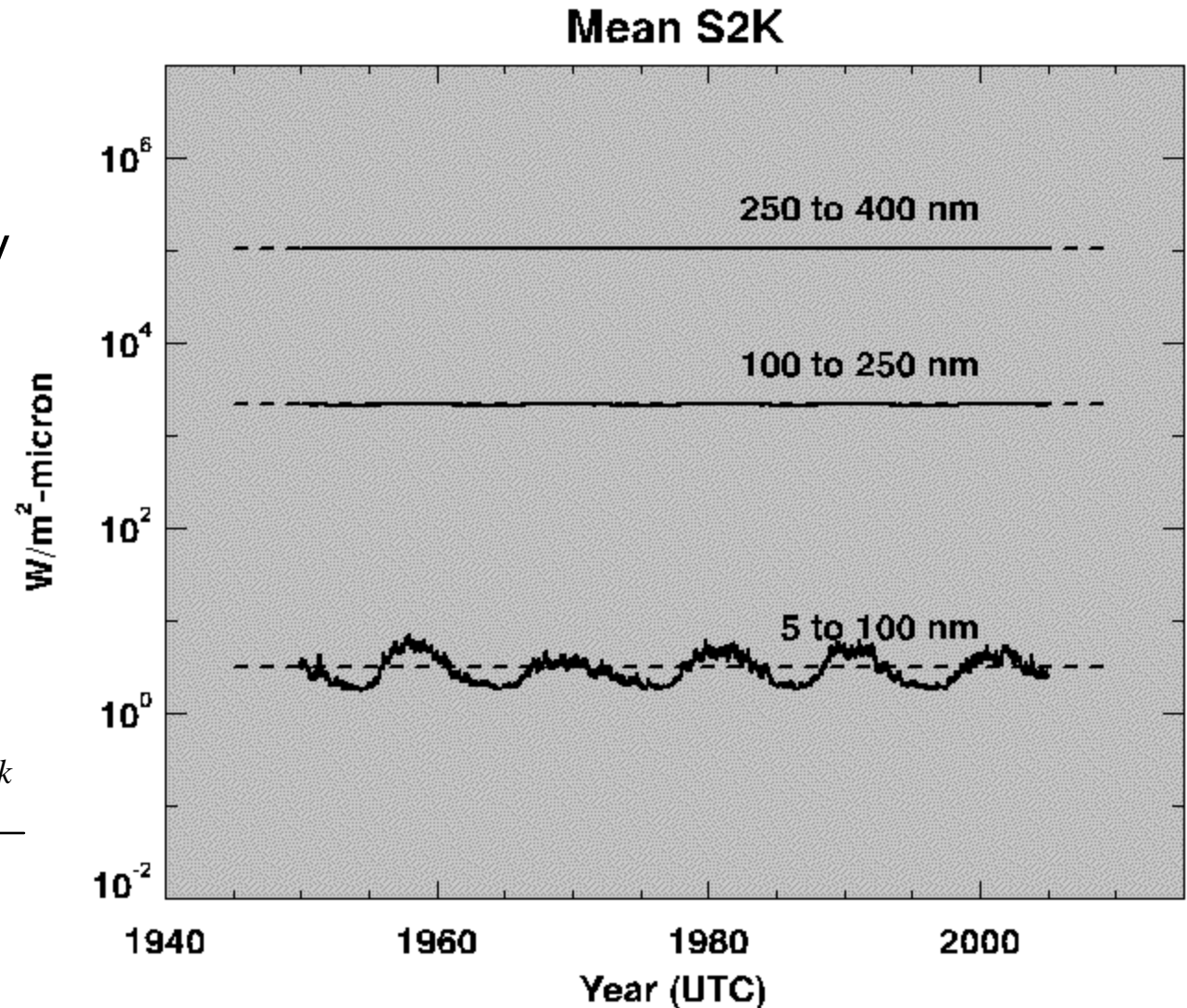
## Solar2000

- Static ASTM E490
- Variable XUV/EUV

## Reference Spectrum

- Mean S2K

$$I_{\lambda_a - \lambda_b} = \frac{\sum_{k=1}^n I(\lambda_k) d\lambda_k}{\sum_{k=1}^n d\lambda_k}$$

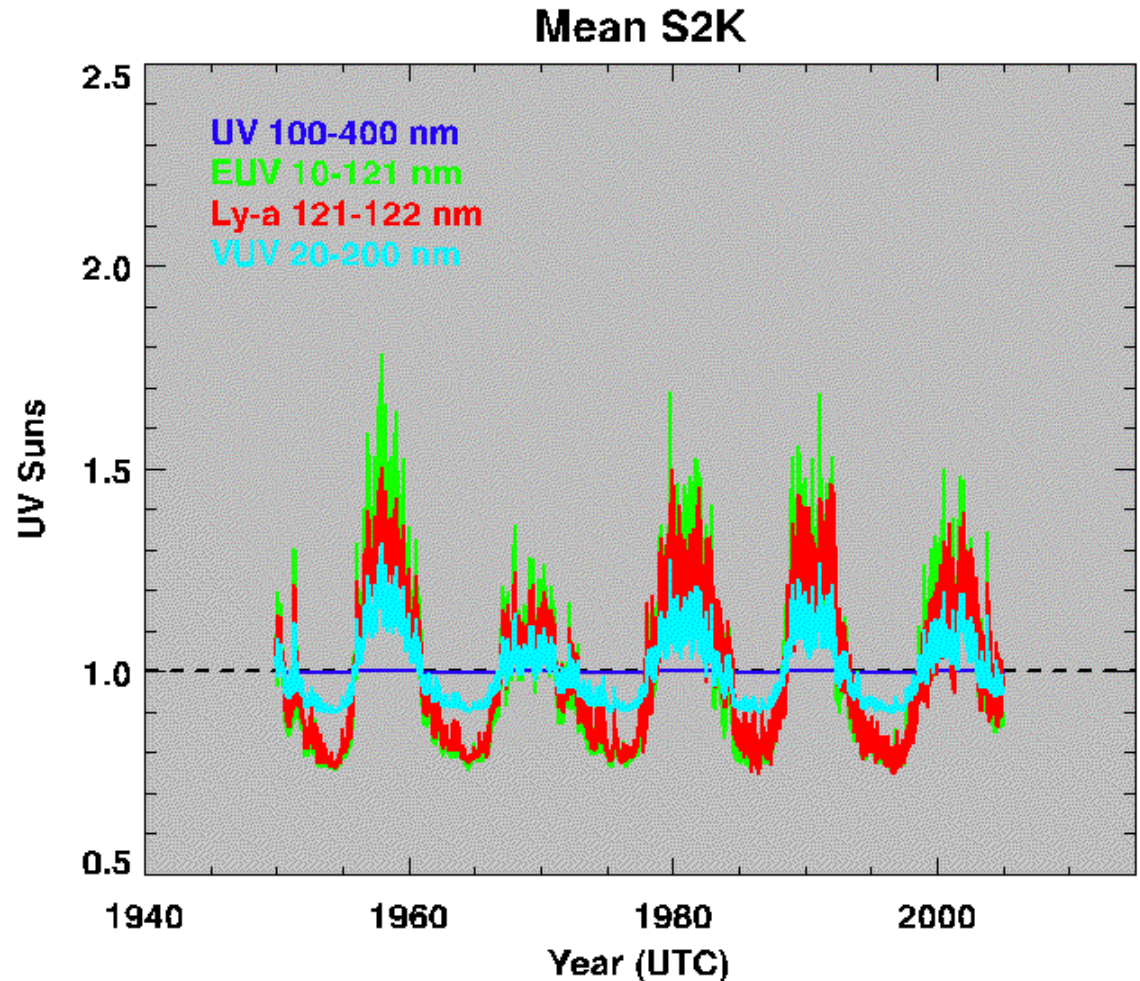






# Solar UV/VUV Variability

- Solar intensity in terms of UV Suns based on mean S2K design model
- Solar source yields UV Suns exceeding unity when mean model is used as reference spectrum



$$UV\ Suns_{\lambda_a-\lambda_b} = \frac{I_{\lambda_a-\lambda_b,source}}{I_{\lambda_a-\lambda_b,reference}}$$



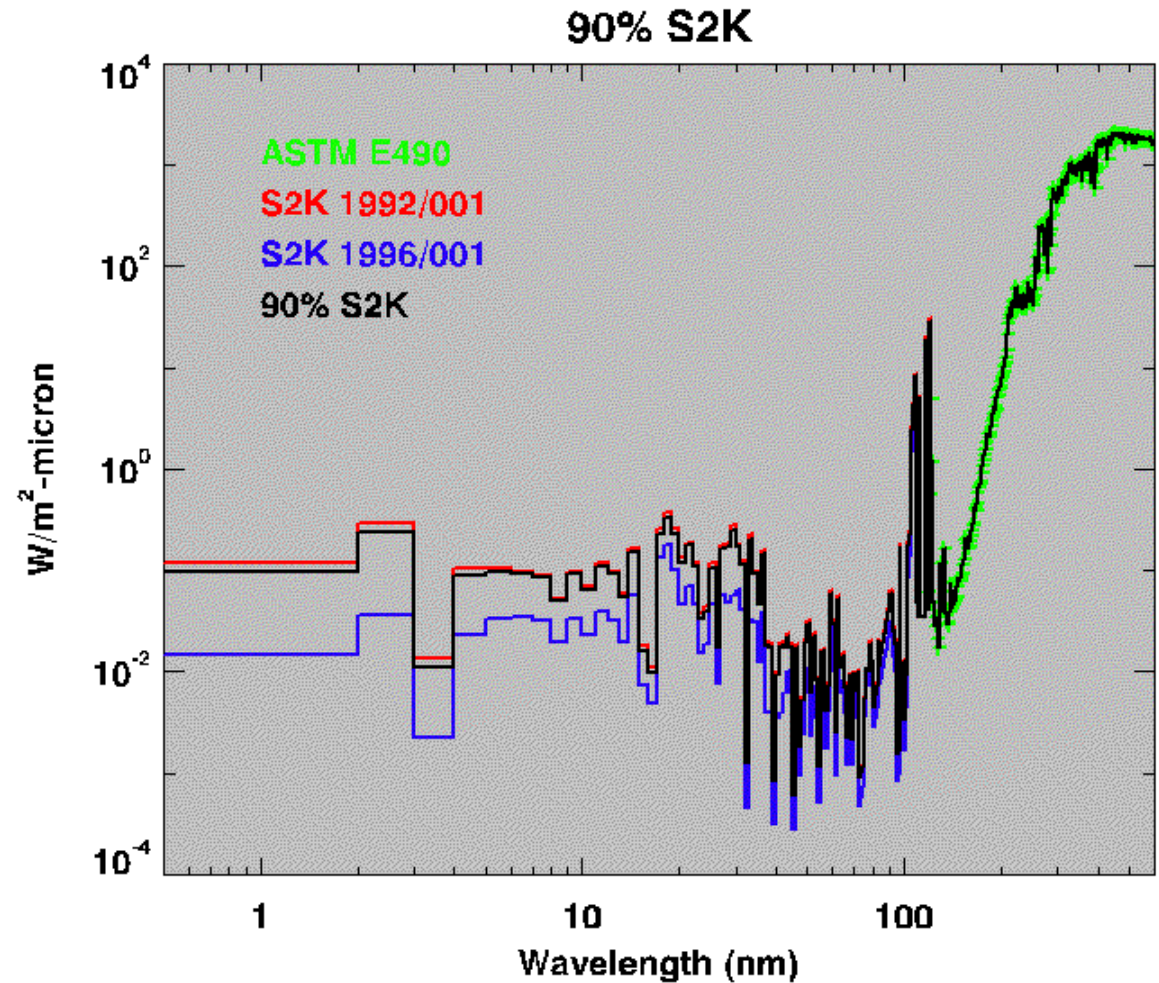
# Solar UV/VUV Variability

## Solar2000

- Static ASTM E490
- Variable XUV/EUV

## Reference Spectrum

- 90% S2K







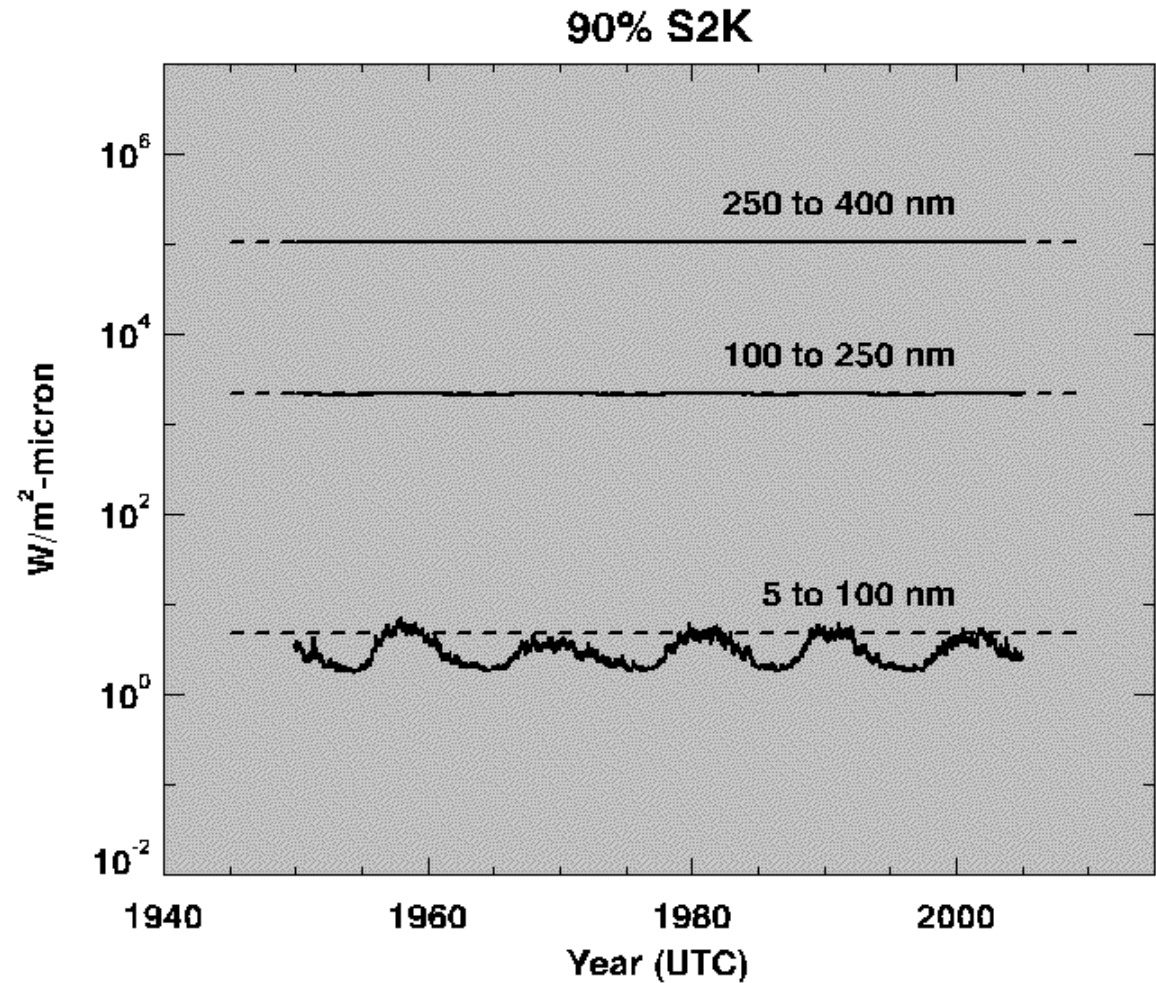
# Solar XUV/EUV/UV Variability

Solar2000

- Static ASTM E490
- Variable XUV/EUV

Reference Spectrum

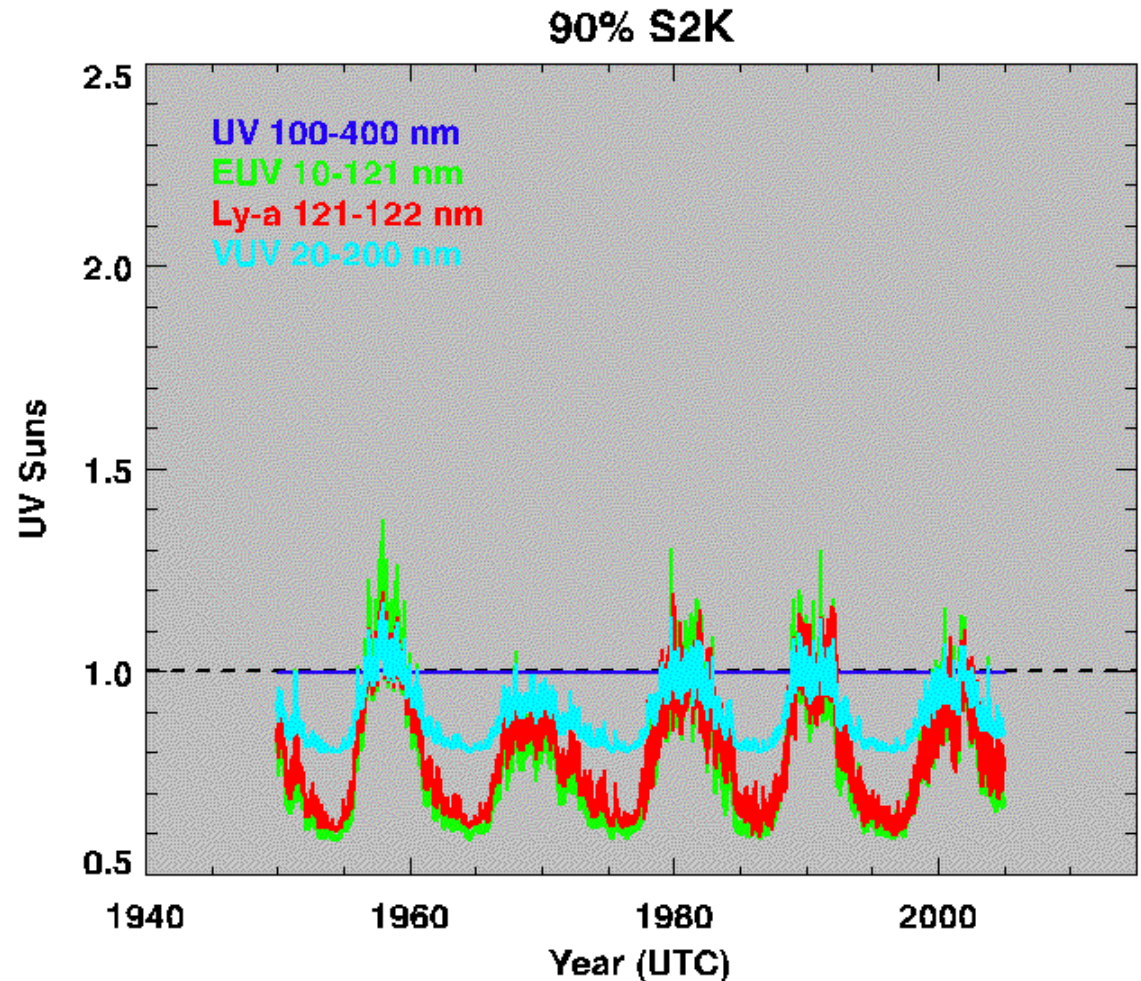
- 90% S2K





# Solar UV/VUV Variability

- Solar intensity in terms of UV Suns based on 90% S2K design model
- Conservative model yields fewer UV sun values exceeding design models

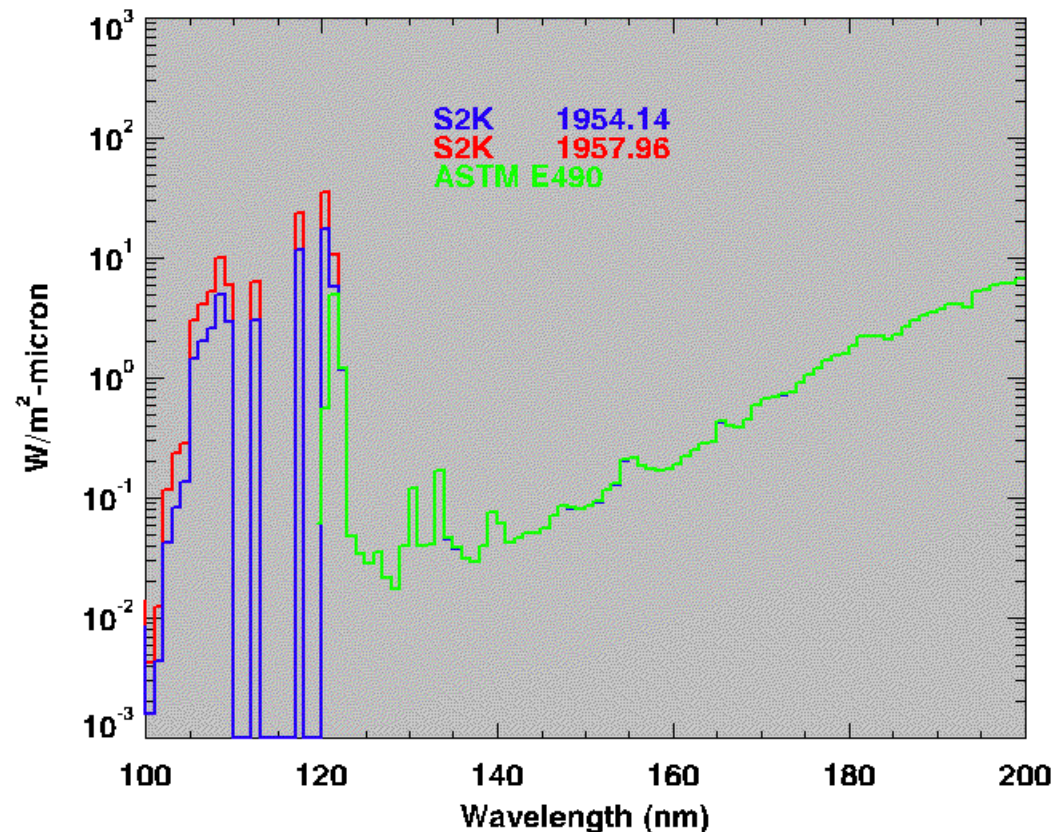






# ASTM E490 and Solar Variability: Spectrum

- S2K model intensity exceeds ASTM-E490 at Lyman- $\alpha$  wavelengths
- S2K as ASTM-E490 for wavelengths longer than Lyman- $\alpha$
- Materials sensitive to wavelengths shorter than Lyman- $\alpha$  may under perform in space environment if
  - qualified only to ASTM-E490
  - degradation dominated by <200 nm environment





# ASTM E490 and Solar Variability: Time

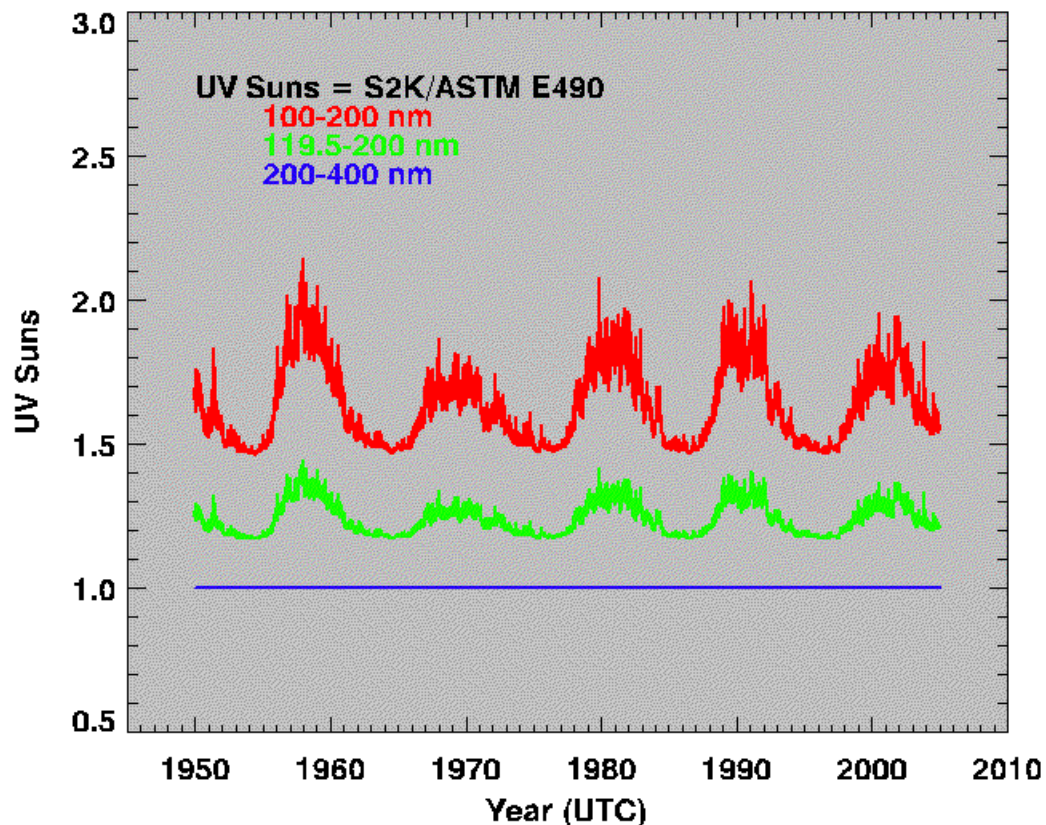
- UV Suns based on Solar2000 exceed the ASTM-E490 environments for all solar cycles from 1950 through 2004 for wavelengths shorter than Lyman- $\alpha$

- ASTM under represents the solar spectrum?

UV source: UARS

- Solar2000 XUV/EUV data measured on orbit [*Tobiska and Bouwer, 2006*]:

|       |        |
|-------|--------|
| TIMED | AE-E   |
| SOHO  | SOLRAD |
| SORCE | YOHKOH |
| SNOE  |        |

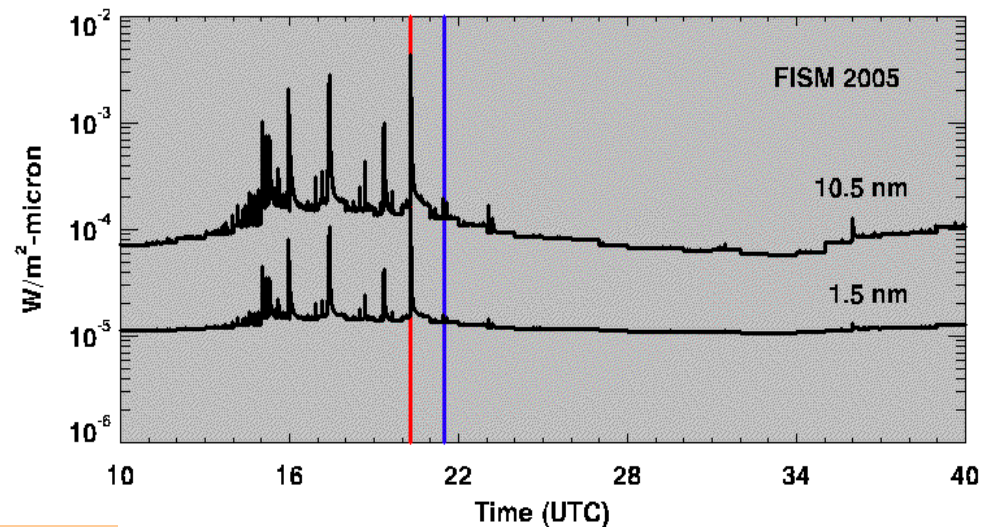
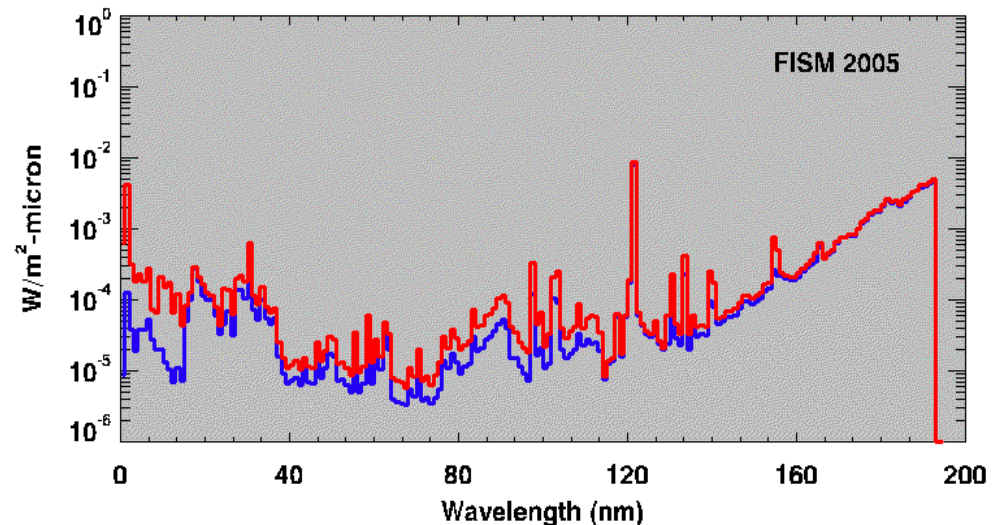






# Flare Irradiance Spectral Model (FISM)

- Empirical solar irradiance model developed by LASP/CU [Chamberlin et al., 2007]
  - Resolution  
 $\Delta\lambda = 1 \text{ nm}$     $0.1 \text{ nm} < \lambda < 194 \text{ nm}$   
 $\Delta t = 60 \text{ seconds}$
  - Data sources
    - Solar Extreme Ultraviolet Experiment (SEE)/TIMED
    - Solar Stellar Irradiance Comparison Experiment (SOLSTICE)/UARS
- FISM developed to provide VUV solar spectral irradiances for input to ionosphere, thermosphere models
  - 100% coverage from 1986 to present



FISM URL:  
<http://lasp.colorado.edu/LISIRD/fism.htm>

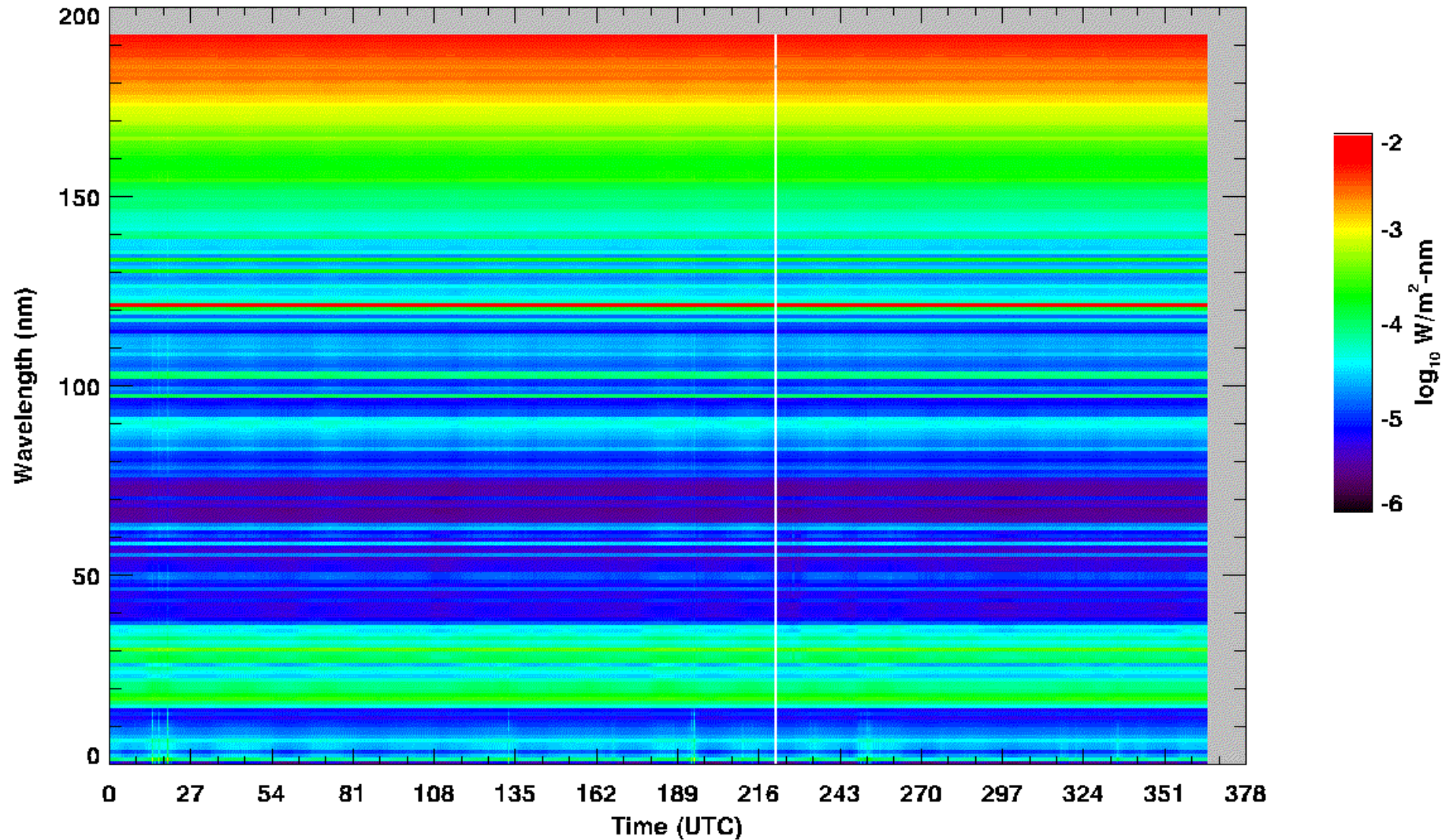




# Solar VUV 2005

## Solar minimum

FISM 2005

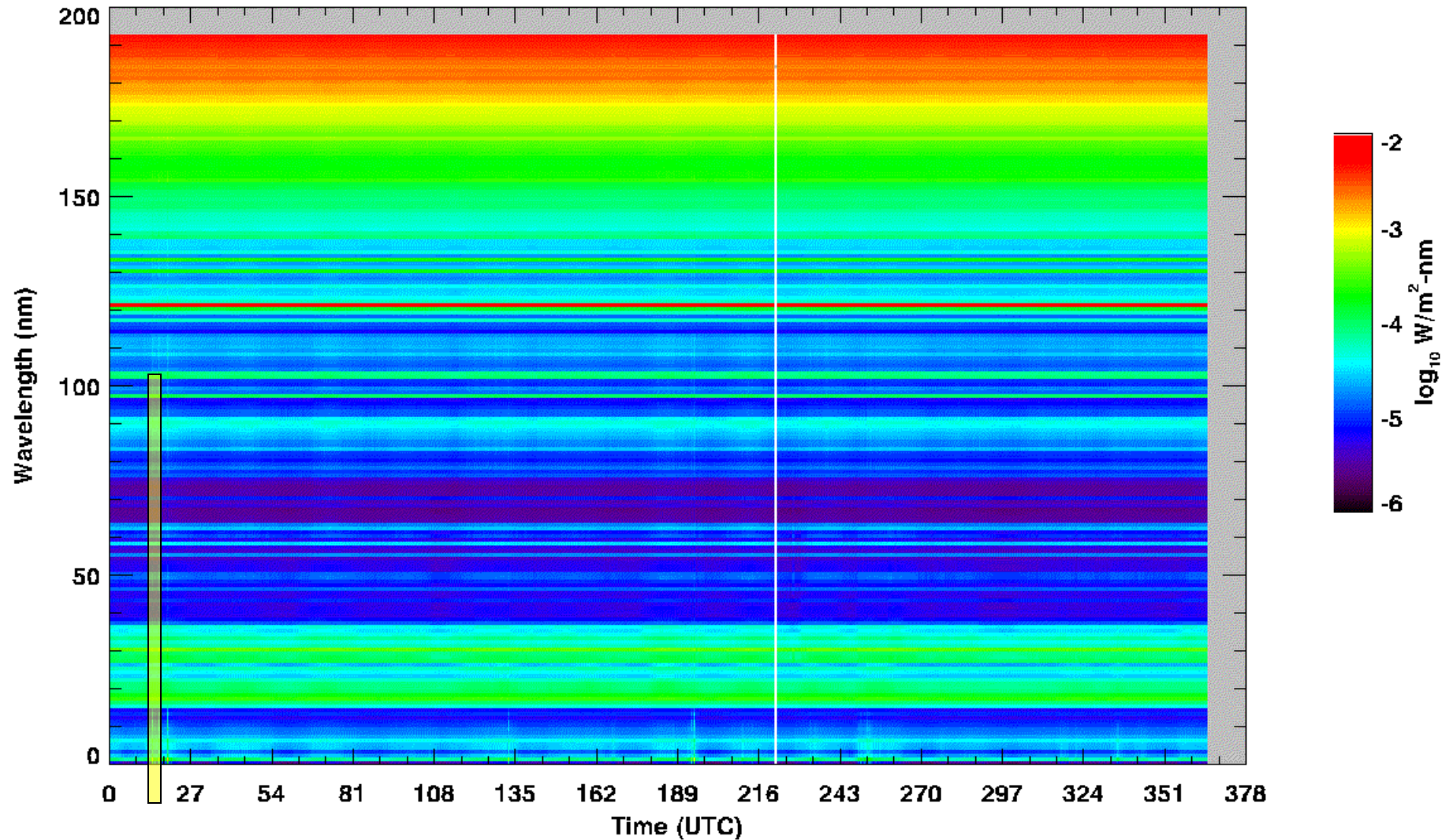




# Solar VUV 2005

## Solar minimum

FISM 2005

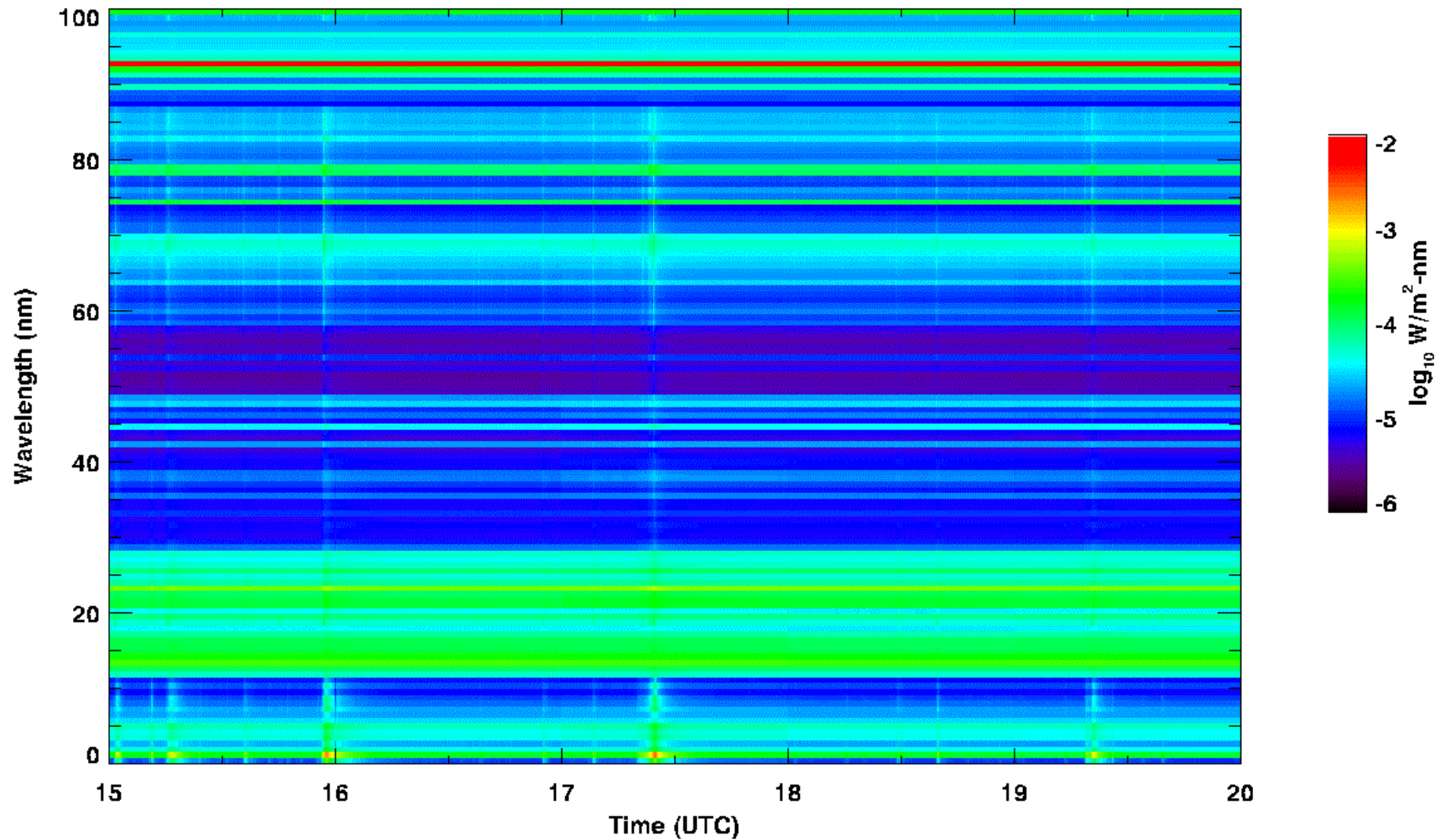




# Solar VUV 2005

## Solar minimum

FISM 2005



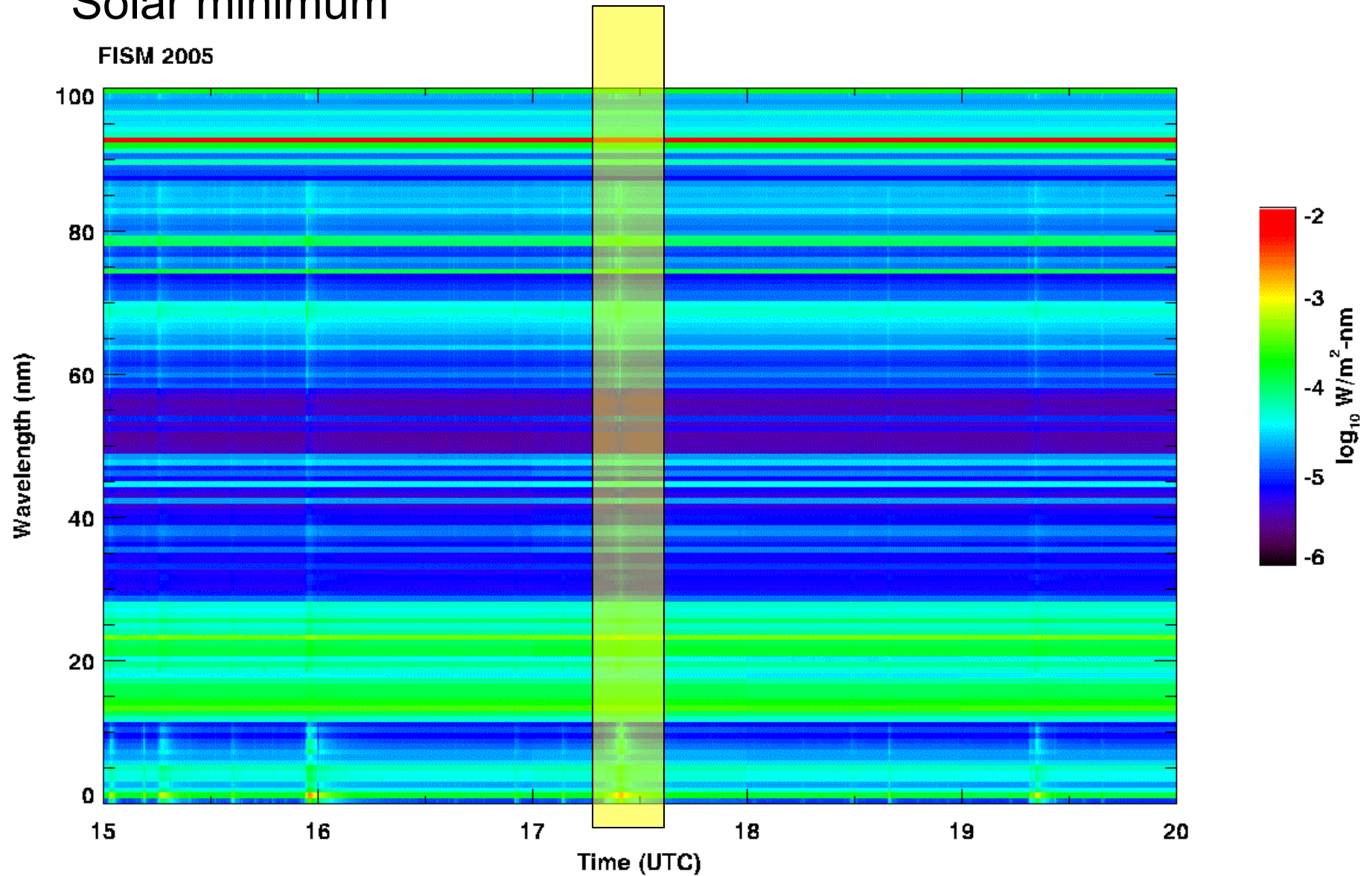




# Solar VUV 2005

## Solar minimum

FISM 2005

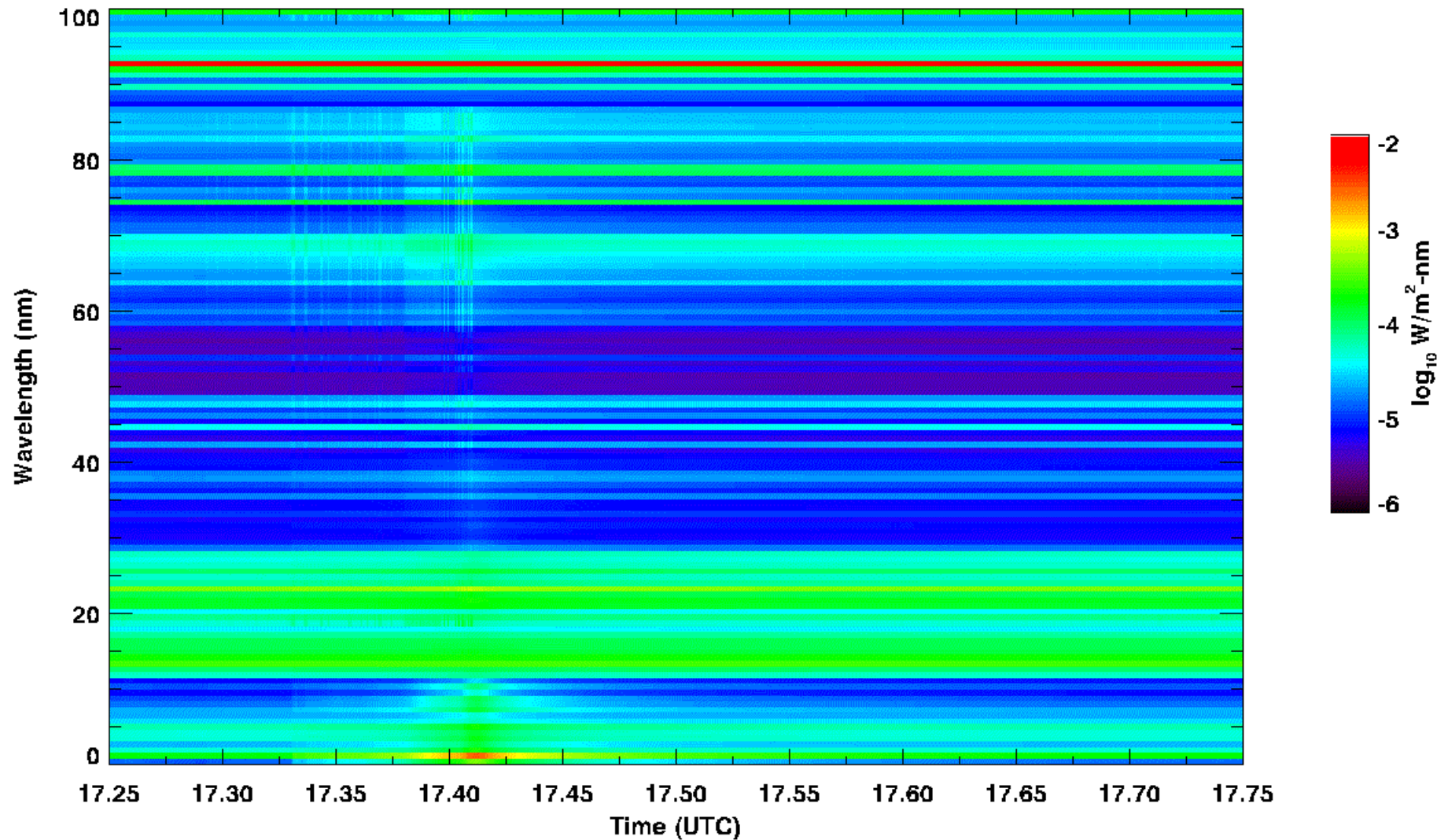




# Solar VUV 2005

## Solar minimum

FISM 2005







# Summary

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- **Static design models typically used to establish laboratory test protocols for determining material response to the UV/VUV space environment**
  - Need to be careful with choice of reference model if test results are to be used for qualifying materials for extended use in space
- **Space climatology and space weather models provide a useful technique for evaluating projected on orbit performance to a “static” design specification**
- **Solar2000 and FISM models are useful tools for**
  - Characterizing dynamic changes in on-orbit solar VUV environments
  - Developing appropriate design environments for screening materials to VUV environments